#### ASTRONOMICHESKII TSIRKULYAR

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**Abstract.** This text contains Instructions for our Authors; at the same time, it shows the general layout of a contribution to the Astronomicheskii Tsirkulyar as it will appear published in the electronic form or printed on paper. See the source file 'sample\_e.tex' stored in this directory to get an idea on how we got this text. Place your abstract here, it must not exceed ten lines. It is in English, if the article is in English; in this case, place the Russian abstract at the end of the article.

#### Introduction

The text of your article starts here.

We accept articles in all branches of astronomy. Preference is given to short notes containing information that should be distributed as soon as possible (e.g., discoveries of supernovae, novae, comets; important novel theoretical results, etc.). Articles submitted should be accompanied by a recommendation from a seminar specialized in the given branch of astronomy; please justify in brief the necessity of rapid publication. All submitted papers are sent to referees — experts in the field of astronomy.

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# Formulas, Illustrations, and Tables

Below we give you some examples of how to compose equations and tables and to insert figures.

Evident inequalities:

$$\pi \gtrsim 3.141$$
 and  $\pi \lesssim 3.142$ 

(this formula is not numbered).

Here is the Saha equation

$$n_e \frac{n^+}{n_1} = \frac{g^+}{g_1} \frac{2(2\pi mkT)^{3/2}}{h^3} \exp\left(-\frac{\chi_1}{kT}\right),$$
 (1)

where  $n^+$  is the number density of ionized atoms in the ground state (cm<sup>-3</sup>),  $g^+$  is the statistical weight of this state,  $n_e$  is the number density of free electrons (cm<sup>-3</sup>).

An example of an illustration (EPS format is preferable):

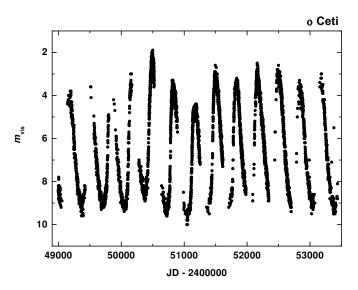


Figure 1: Plot of stellar magnitudes versus time for the star o Ceti.

Equation of radiative transfer in spherical coordinates:

$$\cos \vartheta \frac{\partial I}{\partial r} - \frac{\sin \vartheta}{r} \frac{\partial I}{\partial \vartheta} = -\alpha I + \varepsilon, \tag{2}$$

where  $\varepsilon$  is volume emissivity.

In cylindrical coordinates, the gravitational potential of a body with an arbitrary distribution of density  $\sigma$  is:

$$U(r,\varphi,z) = G \iiint_{V} \frac{\sigma(\rho,\psi,\zeta)\rho \,d\rho \,d\psi \,d\zeta}{\sqrt{\rho^2 + r^2 - 2\rho r \cos(\varphi - \psi) + (z - \zeta)^2}},\tag{3}$$

where the integration is carried out over the entire volume V of the body.

J2000 Other Names Star Name  $m_V$ BD, CPD HD $\delta$ (mag)  $\alpha$ 06<sup>h</sup>45<sup>m</sup>10<sup>s</sup>762  $\alpha$  CMa Sirius  $-16^{\circ}1591$ 48915  $-16^{\circ}41'57''.82$ -1.4606 23 57.005  $\alpha$  Car Canopus  $-52^{\circ} 914$ 45348  $-52\ 41\ 45.55$ -0.57 $+19^{\circ}2777$ 14 15 43.458  $+19\ 12\ 36.73$ -0.05 $\alpha$  Boo Arcturus 124897  $\alpha$  Lyr Vega  $+38^{\circ}3238$ 172167 18 36 55.377  $+38\ 46\ 46.78$ +0.03

Table 1: An example of a small table. Four brightest stars

Notes to Table 1. Here you can insert some comments to the above Table.

## Citations

References in the text should be given in square brackets [1, 2]. Use the abbreviations accepted now in astronomy journals (see, e.g., the Astrophysical Journal):

AZh for Astronomicheskii Zhurnal

PAZh for Pis'ma v Astronomicheskii Zhurnal

ApJ for Astrophysical Journal

AJ for Astronomical Journal

A&A for Astronomy and Astrophysics

MNRAS for Monthly Notices of the Royal Astronomical Society etc.

#### References

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### Название статьи на русском языке

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**Резюме.** В конце статьи, написанной по-английски, приводится её резюме на русском языке.