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except for Prepositions and Articles**A.A. Author1¹, B.B. Author2², and C.C. Authors3³**¹*Author1's address and affiliation**E-mail: Author1's e-mail*²*Author1's address and affiliation*³*Author1's address and affiliation*

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Abstract. This text contains Instructions for our Authors; at the same time, it shows the general layout of a contribution to the *Astronomicheskii Tsirkulyar* as it will appear published in the electronic form. See the source T_EX file ‘sample_e.tex’ stored in this directory to get an idea on how we got this text. Place your abstract here, it must not exceed ten lines. It is in English, if the article is in English; in this case, place the Russian abstract at the end of the article.

Introduction

The text of your article starts here.

We accept articles in all branches of astronomy. Short notes requiring urgent publication (discoveries of supernovae, novae, comets; important novel theoretical results) are preferred. Articles submitted should be accompanied by a recommendation from a seminar specialized in the given branch of astronomy; please justify in brief the necessity of rapid publication. All submitted papers are sent to referees — experts in the field of astronomy.

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Formulas, Illustrations, and Tables

Below we give you some examples of how to compose equations and tables and to insert figures.

Evident inequalities:

$$\pi \gtrsim 3.141 \text{ and } \pi \lesssim 3.142$$

(this formula is not numbered).

Here is the Saha equation

$$n_e \frac{n^+}{n_1} = \frac{g^+}{g_1} \frac{2(2\pi mkT)^{3/2}}{h^3} \exp\left(-\frac{\chi_1}{kT}\right), \quad (1)$$

where n^+ is the number density of ionized atoms in the ground state (cm^{-3}), g^+ is the statistical weight of this state, n_e is the number density of free electrons (cm^{-3}).

An example of an illustration (*EPS format is preferable*):

Equation of radiative transfer in spherical coordinates:

$$\cos \vartheta \frac{\partial I}{\partial r} - \frac{\sin \vartheta}{r} \frac{\partial I}{\partial \vartheta} = -\alpha I + \varepsilon, \quad (2)$$

where ε is volume emissivity.

In cylindrical coordinates, the gravitational potential of a body with an arbitrary distribution of density σ is:

$$U(r, \varphi, z) = G \iiint_V \frac{\sigma(\rho, \psi, \zeta) \rho d\rho d\psi d\zeta}{\sqrt{\rho^2 + r^2 - 2\rho r \cos(\varphi - \psi) + (z - \zeta)^2}}, \quad (3)$$

where the integration is carried out over the entire volume V of the body.

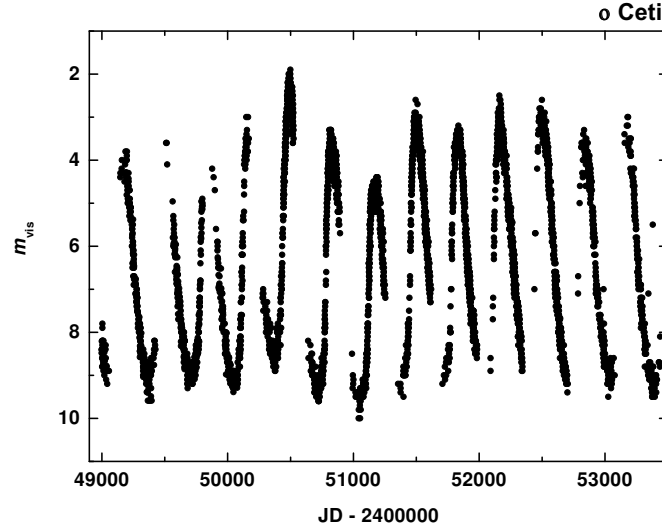


Figure 1: Plot of stellar magnitudes *versus* time for the star *o* Ceti.

Table 1: An example of a small table. Four brightest stars

Star	Name	Other Names		J2000		m_V (mag)
		BD, CPD	HD	α	δ	
α CMa	Sirius	$-16^\circ 1591$	48915	$06^h 45^m 10^s.762$	$-16^\circ 41' 57''.82$	$-1^m.46$
α Car	Canopus	$-52^\circ 914$	45348	06 23 57.005	$-52\ 41\ 45.55$	-0.57
α Boo	Arcturus	$+19^\circ 2777$	124897	14 15 43.458	$+19\ 12\ 36.73$	-0.05
α Lyr	Vega	$+38^\circ 3238$	172167	18 36 55.377	$+38\ 46\ 46.78$	$+0.03$

Notes to Table 1. Here you can insert some comments to the above Table.

Citations

References in the text should be given in square brackets [1, 2]. Use the abbreviations accepted now in astronomy journals (see, e.g., the *Astrophysical Journal*):

AZh for *Astronomicheskii Zhurnal*

PAZh for *Pis'ma v Astronomicheskii Zhurnal*

ApJ for *Astrophysical Journal*

AJ for *Astronomical Journal*

A&A for *Astronomy and Astrophysics*

MNRAS for *Monthly Notices of the Royal Astronomical Society*

etc.

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Резюме. В конце статьи, написанной по-английски, приводится её резюме на русском языке.